

ROCKEFELLER SEES 'CIRCUS FROM LUNA

Gleeful as He Views Antics of Performers and Buys Red Lemonade.

DIMES FOR EVERYBODY

John D. Jr., and Family Also Enjoy Clowns and Peanuts at Pocantico Hills.

John D. Rockefeller renewed his youth yesterday afternoon when he sat under a circus tent at Pocantico Hills and watched a special performance staged from Luna Park.

He shook his head when the juggler kept producing water from a pitcher after he had let half a dozen boys put their hand inside of it to see that it was empty.

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"Two," the man replied, and Mr. Rockefeller laid two dimes on his palm. "Father and mother living?" The man nodded his head and two more dimes were forthcoming.

This incident attracted a crowd and Mrs. Rockefeller was through he had given out \$20 worth of dimes to every member of the troupe and the band.

"I hope you live to be a century," Mr. Rockefeller said one clown, and then he shook hands. Mr. Rockefeller smiled and said: "You people give so much pleasure."

Then he met the contortionist. "How much do you weigh?" "One hundred and thirty-five pounds," replied the man.

Mr. Rockefeller marveled and proceeded to give him a dime for every member of the family.

Just then John D. Rockefeller, Jr., remarked: "Father, there are some on the other side."

The older Rockefeller turned and saw the fat woman and her nurse, to whom he gave dimes. Then he patronized the booths in the midway, buying popcorn, lemonade, peanuts and everything that goes with a circus.

"Who wants to ride on the Rockefeller pony?" yelled the Rev. Werth, who acted as Barker for Winthrop and Lawrence. "Ten cents a ride."

The ponies did a fine business. The circus was given for the benefit of the Community Church, which cost \$1,000.

Mr. Rockefeller and some neighbors gave \$50,000, leaving \$5,000 to be raised by the people. Mr. Rockefeller offered to match dollar for dollar, and in spite of the bad weather, it was hoped that a large sum would be realized.

IN PRIESTHOOD 26 YEARS.

Rev. Fathers Mueller and Werth Celebrate Event to-day.

Twenty-six years ago to-day the Very Rev. Benedict Mueller, provincial of the Province of St. Joseph of the Capuchin Fathers, and the Rev. Ludwig Werth, superior of the Roman Catholic Church of St. John the Baptist, West Thirtieth street, and spiritual director of the Third Order Conference, were ordained to the priesthood by the late Archbishop Francis X. Katz of Milwaukee.

Both will to-day observe that event. Father Werth's parishioners assisting in the observance.

Father Mueller became the provincial of St. Joseph's Province, with headquarters at Detroit, after he had taught at St. Lawrence's College, Mount Calvary, Wis., for twenty-two years, while his classmate was assigned to this city.

2,000 POLICEMEN ELIGIBLE.

A list of names of 2,000 patrolmen who passed the examinations held August 6, 1921, to fill vacancies in the grade of sergeant was made public yesterday.

There are now fifty-six vacancies, but they cannot be filled until Commissioner Enright returns from abroad.

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TINY, DIM STAR DISCLOSED AS TWINS BIGGER THAN SUN

Continued from First Page.

tra were photographed as frequently as possible and the positions of the doubled lines were seen to be continually changing. Sometimes, the stronger lines were displaced to the red and the weaker to the violet end of the spectrum and sometimes the opposite was the case. The former indicated that the brighter star was receding, the fainter approaching us and vice versa.

Further spectra showed that the brighter star was revolving at a speed of 128 miles and the fainter at 154 miles a second and that it took nearly fourteen and a half days to make a complete revolution around each other. A comparatively simple calculation then shows that the two bodies must be at a distance of about 55,000,000 miles apart, three-fifths the distance from the earth to the sun and slightly less than the distance of Venus from the Sun. If the two stars combined were of the same mass as Venus and the Sun they should revolve in about the same period, 225 days, but as a matter of fact, they revolve in fourteen days, one-fifth of the time and at a speed over seven times as great as Venus. This evidently indicates a very much stronger gravitational pull between the two bodies and a very much greater mass.

If we know the period of revolution and the separation of two bodies in space, two suns, or a sun and planet, or a planet and satellite, we can easily calculate the mass of the two by a comparatively simple relation derived from the law of gravitation called the Harmonic Law. This law states that the combined mass of the revolving system is proportional to the cube of the separation divided by the square of the period. In all these calculations we use the Earth Sun system as the standard and as the Earth has only one three hundred and thirty thousandths the mass of the Sun, it can be neglected and we get the mass of the required system in terms of the Sun's mass.

Hence in this case we have the double star revolving in a period of 14.4 divided by 365.25, about one-twenty-fifth of the Earth's period, the year, at a distance of 55,000,000 divided by 93,000,000, or about three-fifths the Sun's distance. From the Harmonic Law then the mass of this double star will be three-fifths cubed, divided by one-twenty-fifth squared the mass of the Sun, which comes out 135 times the Sun's mass. If we use the exact fractions we find that this system has 138.9 times the mass of the Sun, that the brighter star is 75.6 and the fainter 63.3 times the Sun's mass, or about four times greater than for any previously discovered system.

A further calculation from the spectroscopic observations shows that the two bodies are revolving in an ellipse around one another of nearly the same shape as the orbits of the planets. But the spectroscopic results cannot tell us anything of the inclination of the planet of this ellipse to the line of sight. If this plane passes through the earth, the distance and masses are as given above, but if not, they must be greater than calculated.

From considerations of the probable size of the two stars and their distances apart, it is practically certain that this plane is inclined more than fifteen degrees, else the two bodies would mutually eclipse each revolution and there would be a decrease in the line which has not been observed. This condition makes it probable the combined mass is 160 times the Sun, while if we consider all inclinations equally likely the probable average value would be 220 times the Sun's mass.

May Revise Stellar Theories.

The very great mass of these two bodies shows that they must be exceptional in many respects and this discovery may require some revision of the theories of the formation and constitution of the stars. However that may be, we can easily deduce further interesting data about these two tremendous stars. I have already said that they are about forty times as bright as the Sun for each unit of area and as they are so much more massive they must be much bigger.

Two methods are available for determining how much brighter they are than the Sun. The first is wholly theoretical, developed by Prof. Eddington of Cambridge, from which we find that the brighter star is about 29,000 and the fainter 15,000 times as bright as the Sun.

However, there is some uncertainty about these values and we can use another method depending upon the surface brightness and the area of the stars. It is probable from various considerations that such excessively hot stars are very much less dense than the Sun, probably only one hundredth as dense. As the brighter star has at least seventy-five times the mass and one-hundredth the density, it must have 7,500 times the volume, or 19.5 times the diameter of the Sun. Its area will then be 19.5 squared times or 380 times the Sun and if as already stated this relativity is forty times as bright, it will be more than 15,000 times as bright as the Sun.

It is admitted by Eddington that his computed values are somewhat too high for the very hot stars, so that we may take as a moderate estimate of the brightness that the brighter star is 15,000 and the fainter 12,000 times as bright as the Sun, much brighter than any known star in the sky. The second brightest known star is Rigel, which is, however, only half the brightness of this pair. Knowing the actual brightness 27,000 times the Sun, the apparent brightness, 6.1 magnitude and the fact that the Sun would appear of the same brightness as this star if it were removed to a distance of sixty light years, makes the actual distance of this star sixty times the square root of 27,000, or about 10,000 light years.

This double star system is moving away from the Sun at the rate of fifteen miles a second, but there is also a cloud of calcium vapor either surrounding or else between the two and the pair. The spectroscopic shows this vapor to be moving away from us at a constant speed of ten miles per second. If this cloud surrounds the system, it should move with the same speed and a possible explanation of the discrepancy is the relativity shift of the spectrum lines. Einstein's theory calls for three astronomical consequences: the advance of the perihelion of Mercury which is fully confirmed, the bending of light passing near the Sun, which has been observed at the 1919 eclipse, and is to be again observed in Australia on September 21, and a displacement of the lines of the spectrum of a heavy body to the red. This last effect for the Sun is equivalent to a speed of slightly less than half a mile per second away from us, but it has not yet been definitely found in the Sun. As the shift is directly proportionate to the mass and diminishes with the diameter of the star, this star should be only seven instead of twenty times that of the Sun and this would require the unlikely great density of four-tenths the Sun instead of a probable density one hundredth.

under our little sun, which, by the by, is not cutting so much of a swath among the two million stars that are being brought closer daily to our vision. Every day we look for the discovery of still larger star suns, and we are getting slowly but steadily to the core of things celestial, even though the stupendous distances we measure in space—ten thousand light years—are incomprehensible to the lay mind.

"Of course we cannot always conceive intelligently to the lay mind the grasp we sometimes get on things infinite. Ten thousand light years is some distance away, and the light from this great double star has taken some time to reach this planet. It started long before man was able to conceive the thought so ably expressed by the psalmist over 3,000 years ago when he sang 'A thousand years in Thy sight is but a yesterday,' but its spectrum recorded on photographic plates tells us more about its existence than we ever expected to have revealed to us in my day.

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"The measurements recorded by Mr. Plaskett must be accepted as most reliable. His constant and constant work photographing its spectra has given us the speed of both the brighter and fainter stars, as well as the exact time taken by both to revolve around each other.

"There is nothing I can add to the intelligible and remarkable report submitted to the Royal Astronomical Society in London by Mr. Plaskett. He has given The New York Herald readers a most comprehensive resume of his investigations and I am sure that its publication will create as great and widespread an interest as the discovery and measurement of Betelgeuse, which was a notable achievement.

"Personally, I have been out of touch with the world for several months through illness, and I am indeed glad on the eve of my recovery to know of the splendid discovery regarding the mass of this great stranger in the distant heavens."

Although stars like Antares and Betelgeuse are much larger in diameter they are extremely tenuous and we have no information about their mass. This star, which is No. 1309 of the six degree zone of Argelander's catalogue, is especially remarkable for its enormous mass, about four times as great as any previously determined, and the discovery and investigation is important chiefly on that account and on its bearing on the theories of the constitution and development of the stars.

PROF. JACOBY PRAISES WORK OF DISCOVERER

Feat Most Interesting Since Betelgeuse Achievement.

"Most remarkable investigation and discovery," said Prof. Harold Jacoby of Columbia University yesterday, "it is outstanding among recent astronomical investigations, and all the more remarkable for emanating from such a reliable investigator and sky explorer as Prof. J. E. Plaskett of Victoria, British Columbia.

"Mr. Plaskett's research work among the stars has been of prime importance and benefit to everybody who spends their nights searching the heavens for new appearances and further knowledge of nebulae, planets and stars, catalogued and uncatalogued. "The great double star, which has revealed through the spectroscopy to Mr. Plaskett its unlooked for properties, has been known to the astronomical world for about seventy years, but it has now suddenly become an extraordinary object to all, not only by reason of the knowledge revealed to us of its elements, but perhaps more so by the fact that its temperature which Mr. Plaskett has discovered to be 30,000 degrees Fahrenheit, placing it among the very hottest if not the most massive of the now known stars. "Of course there is nothing new

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ESCAPES FROM POLICE BY THREAT TO JUMP

Man Resents Wife Having His Sanity Determined.

Patrolman John Mhertens of the West Forty-seventh street police station and Warrant Officer Fisher went to the home of Matthew O'Hare at 417 West Fifty-third street yesterday afternoon to take O'Hare, at the request of his wife, to the West Side court, where she wished to start proceedings to determine his sanity.

Enraged by the presence of the officers in his home, O'Hare stood at bay in a hall window and threatened to jump to the small courtyard below rather than surrender. He smashed one of the windows with an iron bar about two feet long and brandished it in a threatening manner before Patrolman Mhertens drew his gun. Mrs. O'Hare pleaded with the policeman not to shoot, and finally persuaded her husband to climb down from the window and return to their apartment. Later, while she and the officers were waiting in one of the rooms in the

apartment, he went out into the hall and disappeared. He had not returned at a late hour last night.

In August, 1918, O'Hare fell about thirty feet from a scaffolding in one of the yards of the Standard Shipbuilding Corporation and suffered concussion of the brain. He has not been able to work since that time, the slender support of his family coming from compensation from the State Labor Department.

Recently his wife says she has observed symptoms which prompted her to arrange for his examination in Bellevue Hospital. O'Hare is 45 years old, about six feet tall, weighs 175 pounds and has dark complexion, with black hair. When he left home he wore a blue suit and straw hat. He has blue eyes.

NOTE full name. Ludwig Baumann. Open 9 to 5:30—Saturday Evenings until 9:30.

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